

Logitech

Antenna Under Test (AUT)

Report

Model Name: F00015

Equipment Type: PEN

Manufacturer: LOGITECH EUROPE S.A.

Address: EPFL – Quartier de l’Innovation Daniel Borel Innovation
Center, 1015 Lausanne, Switzerland

Tested by: _____ **Alice Künstner, H.Favey** _____

Report Date: _____ **2025/08/21** _____

Report Release History

| Report version | Description | Date Issued |
|----------------|------------------|-------------|
| 1.00 | Original release | 2025/08/21 |

Table of Contents

| | |
|--|----|
| 1. EUT Antenna Information | 3 |
| 2. Measured Values and Calculation of Antenna Gains | 3 |
| 3. Conducted Power Measurement | 4 |
| 3.1 Test Setup | 4 |
| 3.2 Test Instruments | 4 |
| 3.3 Test Procedure | 4 |
| 3.4 Test Result of RF conducted Power | 4 |
| 4. Radiation Pattern Measurement | 5 |
| 4.1 Test Location | 5 |
| 4.2 Description of the anechoic chamber | 5 |
| 4.3 Test Instruments | 6 |
| 4.4 Test Procedure | 7 |
| 4.5 Test Setup | 8 |
| 4.6 2D Pattern Test Plot | 9 |
| Clarifying the Location of Maximum EiRP Measurement | 12 |
| 5.0 Conclusion | 15 |

1. EUT Antenna Information

- 1) Antenna Material : Polyimide printed
- 2) Antenna Type : Dipole
- 3) Antenna Dimension: 17.2 x 10 mm
- 4) Operating Frequency : 2.4 GHz – 2.4835 GHz
- 5) Input Impedance : 50 Ω
- 6) Standing-Wave Ratio : < 2

2. Measured Values and Calculation of Antenna Gains

Measure peak horizontal/vertical EIRP on each x-y, y-z, x-z plane. The highest measured values will be used to calculate the antenna peak gain.

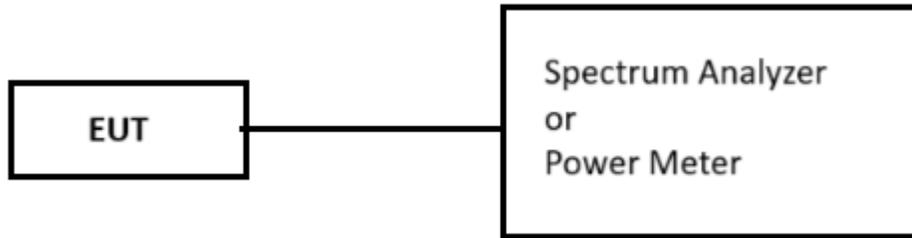
$$\text{Antenna Peak Gain (dBi)} = \text{Max EIRP(dBm)} - \text{Conducted Power (dBm)}$$

| Frequency | X-Y Plane $\phi=0\sim 360^\circ, \theta=90^\circ$ | | X-Z Plane $\phi=0^\circ, \theta=0\sim 360^\circ$ | | Y-Z Plane $\phi=90^\circ, \theta=0\sim 360^\circ$ | | Max Peak EIRP (dBm) | Conducted Power (dBm) | Antenna Peak Gain (dBi) |
|-----------|--|-----------------------|---|-----------------------|--|-----------------------|---------------------|-----------------------|-------------------------|
| | Ver. Peak EIRP (dBm) | Hori. Peak EIRP (dBm) | Ver. Peak EIRP (dBm) | Hori. Peak EIRP (dBm) | Ver. Peak EIRP (dBm) | Hori. Peak EIRP (dBm) | | | |
| 2.402 GHz | -10.6 | -1.3 | -4.4 | 1.6 | 1.1 | -4.7 | 1.6 | 7.97 | -6.37 |
| 2.440 GHz | -9.5 | 0.1 | -2.7 | 2.6 | 2.1 | -4.2 | 2.6 | 7.94 | -5.34 |
| 2.480 GHz | -10.1 | -0.3 | -1.6 | 1.4 | 0.9 | -3.7 | 1.4 | 8.11 | -6.71 |

Test Date: _____ 2025/06/25 _____

3. Conducted Power Measurement

3.1 Test Setup



3.2 Test Instruments

| Description | Model No. | Serial No. | Last Calibration |
|-------------------------------|-----------|------------|------------------|
| Spectrum Analyzer Keysight | N9000B | MY57102570 | 27-JUL-2023 |

Note: The calibration interval of the above test instruments is 24 months

3.3 Test Procedure

A spectrum analyzer or Power meter was used to perform output power measurement, setting the detector to average and configuring EUT continuously transmitting power(100% duty cycle).

3.4 Test Result of RF conducted Power

| Frequency | Conducted Power (dBm) |
|-----------|-----------------------|
| 2.402 GHz | 7.97 |
| 2.440 GHz | 7.94 |
| 2.480 GHz | 8.11 |

Test Date: 2025/06/25

4.3 Test Instruments

| Description | Model No. | Serial No. | Last Calibration |
|--------------------------------------|--|--------------|------------------|
| Anechoic chamber Keysight | Microwave Vision Group: 650Mhz to 18GHz , 4.7 x 4.6 x 3.2 m | N.A. | N.A. |
| Starlab Main Unit | SL V2_065-18GHz | 1102284-0035 | 9 Aug. 2024 |
| Starlab (Active Switching Unit) | | | |
| Starlab (Tx Amplification Unit) | | | |
| Starlab (Transfer Switching Unit) | | | |
| Starlab (Rx Amplification Unit) | | | |
| Spectrum Analyzer | Keysight N9020B MXA Signal Analyzer, Multi- touch, 10 Hz to 26.5 GHz | MY60110720 | 27 Jul 2023 |
| VNA | Agilent E5071C | MY46111439 | 18 Jul. 2024 |
| Radio tester | Rohde & Schwarz CMW500 | 146151 | 26 Aug 2021 |
| Software | WaveStudio for passive and active measurements | V22.1 | |
| Workstation | Asus 841R6 windows 10 | N.A. | N.A. |

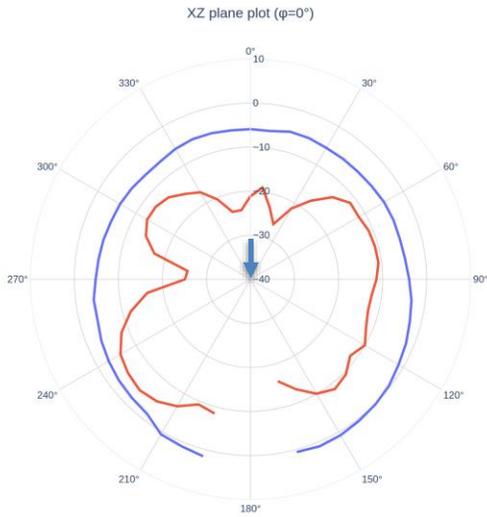
Note: The calibration interval of the above test instruments is 24 and 36 months

4.4 Test Procedure

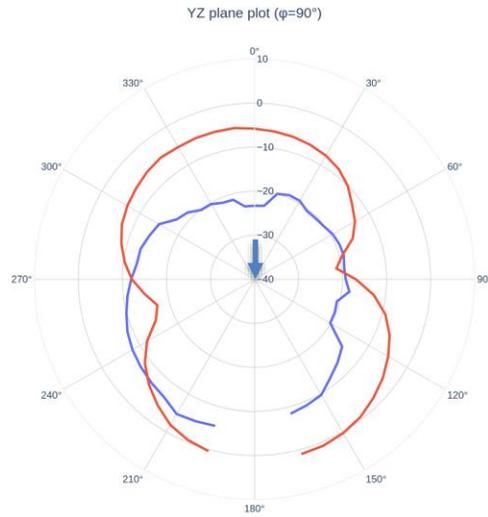
- i. Set the EUT in test mode, fixed frequency, carrier only, 100% duty cycle, at maximum RF power level.
- ii. Place the EUT on the White support (front of the EUT on the X direction) in the Starlab chamber
- iii. Setup the Wave studio software to the frequency set in the EUT (2.402 GHz, 2.440 GHz or 2.480 GHz)
- iv. Set the Starlab chamber to measure the data of emitted radiation power in steps of 10 degrees.
- v. Verify that the signal is stable and not perturbed
- vi. Close the door of the anechoic chamber
- vii. Run the test. The Starlab chamber measures the data of emitted radiation power in steps of 10 degrees. In theta and Phi, for both Horizontal and Vertical antenna polarization.
- viii. Once the test is completed, save the results for both Horizontal and Vertical polarization to a file.
- ix. Record the maximum radiation pattern among the 3 planes XY, XZ, YZ.
Antenna Peak Gain (dBi) = Max EIRP (dBm) - Conducted power (dBm)

4.6 2D Pattern Test Plot

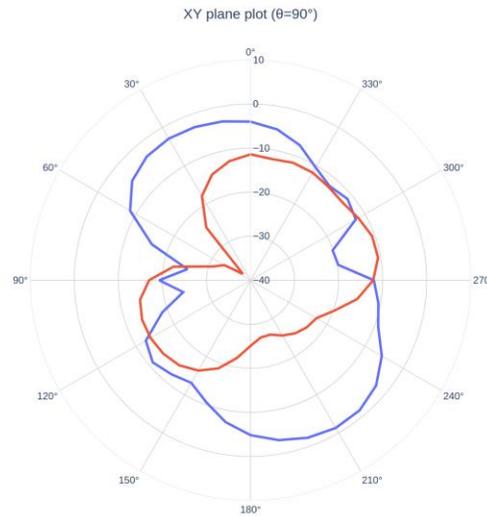
EUT at 2.402 GHz: measured radiated power. Polar 1 = H, Polar 2 = V



— 2402MHz - Polar 1, TRP=-3.51dBm, Pmax=-1.6dBm, neb.h5pmp_001_-_Neb_X888_Tx_Pos8dBm_CH02.tx
 — 2402MHz - Polar 2, TRP=-3.51dBm (0.00dBm), Pmax=-4.4dBm, neb.h5pmp_001_-_Neb_X888_Tx_Pos8dBm_CH02.tx

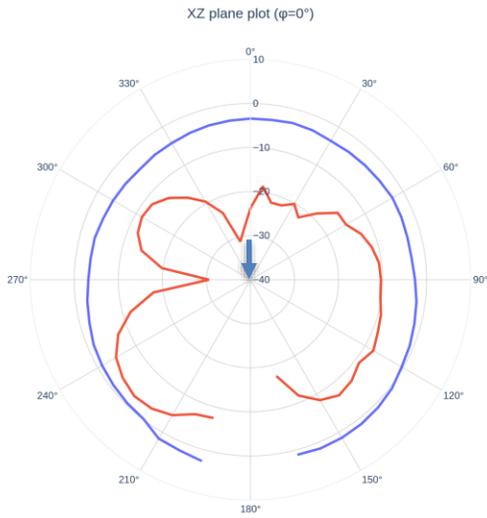


— 2402MHz - Polar 1, TRP=-3.51dBm, Pmax=-4.7dBm, neb.h5pmp_001_-_Neb_X888_Tx_Pos8dBm_CH02.tx
 — 2402MHz - Polar 2, TRP=-3.51dBm (0.00dBm), Pmax=-1.1dBm, neb.h5pmp_001_-_Neb_X888_Tx_Pos8dBm_CH02.tx

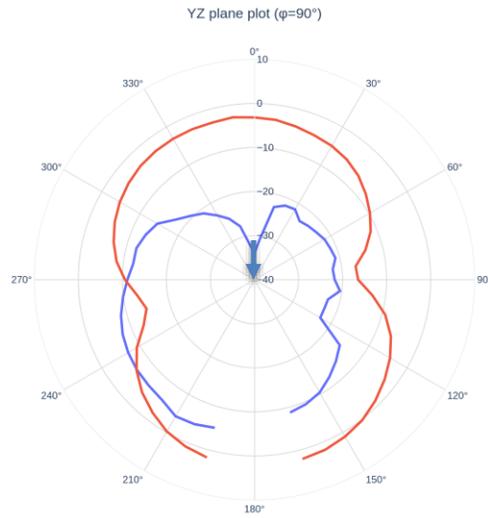


— 2402MHz - Polar 1, TRP=-3.51dBm, Pmax=-1.3dBm, neb.h5pmp_001_-_Neb_X888_Tx_Pos8dBm_CH02.tx
 — 2402MHz - Polar 2, TRP=-3.51dBm (0.00dBm), Pmax=-10.6dBm, neb.h5pmp_001_-_Neb_X888_Tx_Pos8dBm_CH02.tx

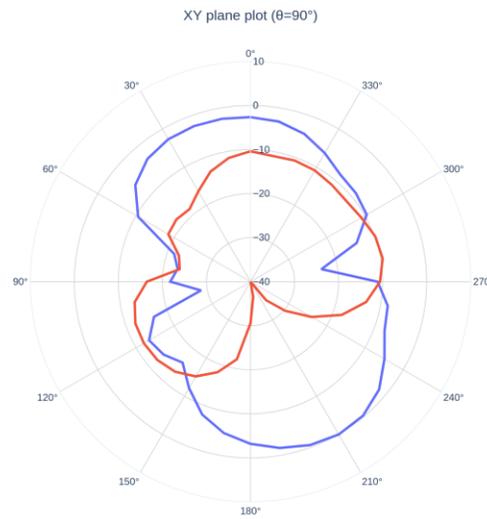
EUT at 2.440 GHz: measured radiated power. Polar 1 = H, Polar 2 = V



— 2440MHz - Polar 1, TRP=-2.27dBm, Pmax=2.6dBm, neb.h5pmp_002_-_Neb_X888_Tx_Pos8dBm_CH40.trx
 — 2440MHz - Polar 2, TRP=-2.27dBm (0.00dBm), Pmax=-2.7dBm, neb.h5pmp_002_-_Neb_X888_Tx_Pos8dBm_CH40.trx

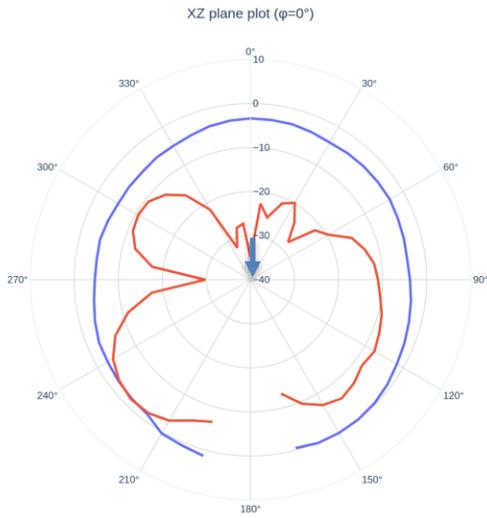


— 2440MHz - Polar 1, TRP=-2.27dBm, Pmax=-4.2dBm, neb.h5pmp_002_-_Neb_X888_Tx_Pos8dBm_CH40.trx
 — 2440MHz - Polar 2, TRP=-2.27dBm (0.00dBm), Pmax=-2.1dBm, neb.h5pmp_002_-_Neb_X888_Tx_Pos8dBm_CH40.trx

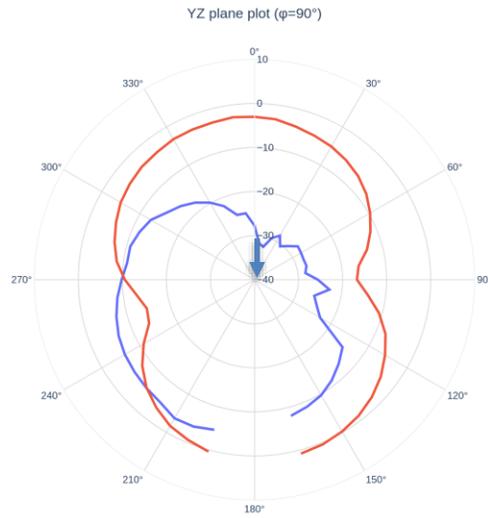


— 2440MHz - Polar 1, TRP=-2.27dBm, Pmax=0.1dBm, neb.h5pmp_002_-_Neb_X888_Tx_Pos8dBm_CH40.trx
 — 2440MHz - Polar 2, TRP=-2.27dBm (0.00dBm), Pmax=-9.5dBm, neb.h5pmp_002_-_Neb_X888_Tx_Pos8dBm_CH40.trx

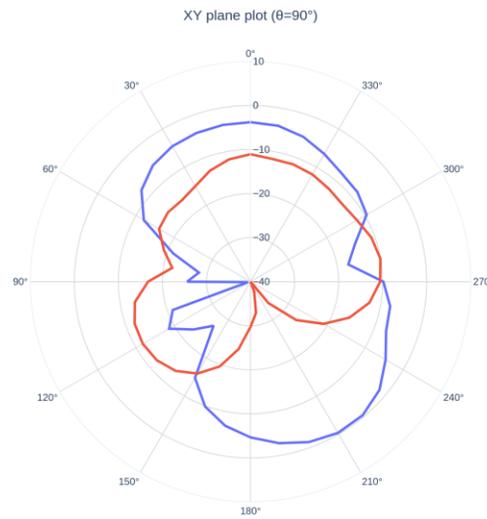
EUT at 2.480 GHz: measured radiated power. Polar 1 = H, Polar 2 = V



— 2480MHz - Polar 1, TRP=-2.94dBm, Pmax=-1.4dBm, neb.h5pmp_003_-_Neb_X888_Tx_Pos8dBm_CH80.trx
 — 2480MHz - Polar 2, TRP=-2.94dBm (0.00dBm), Pmax=-1.6dBm, neb.h5pmp_003_-_Neb_X888_Tx_Pos8dBm_CH80.trx



— 2480MHz - Polar 1, TRP=-2.94dBm, Pmax=-3.7dBm, neb.h5pmp_003_-_Neb_X888_Tx_Pos8dBm_CH80.trx
 — 2480MHz - Polar 2, TRP=-2.94dBm (0.00dBm), Pmax=-0.9dBm, neb.h5pmp_003_-_Neb_X888_Tx_Pos8dBm_CH80.trx



— 2480MHz - Polar 1, TRP=-2.94dBm, Pmax=-0.3dBm, neb.h5pmp_003_-_Neb_X888_Tx_Pos8dBm_CH80.trx
 — 2480MHz - Polar 2, TRP=-2.94dBm (0.00dBm), Pmax=-1.1dBm, neb.h5pmp_003_-_Neb_X888_Tx_Pos8dBm_CH80.trx

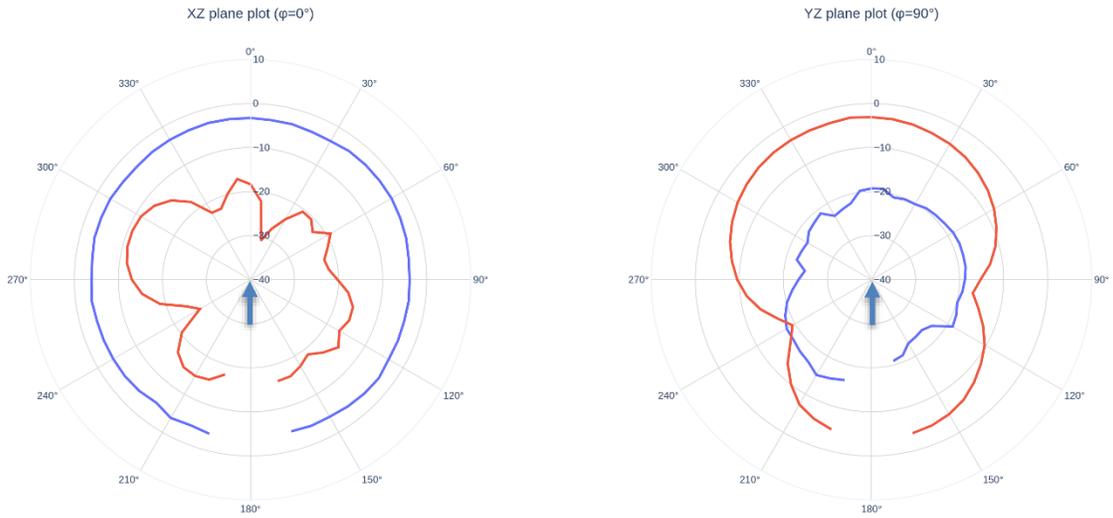
Clarifying the Location of Maximum EiRP Measurement

The Starlab measurement setup includes a blind area, where data cannot be captured, corresponding to the theta angle range of approximately 165° to 195°. To ensure no unintended gain occurs within this blind spot, an additional measurement is conducted by flipping the Device Under Test (DUT) within the Starlab system. This process is indicated by the arrow in the accompanying image, highlighting the actuation direction on the device's switch.

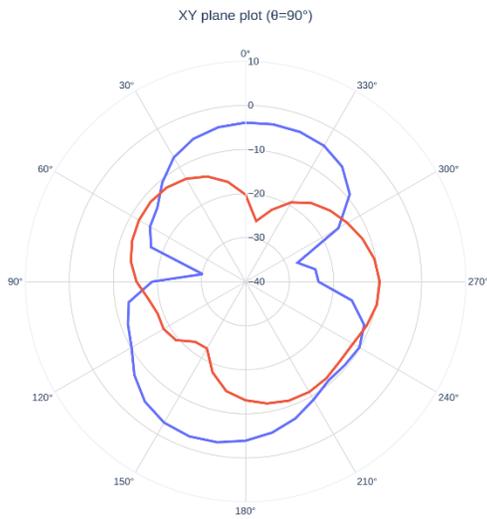
Due to technical constraints, the original DUT could not be reused for this measurement, and a different DUT was employed instead. This explains the observed variation in measurement consistency. Nevertheless, the results demonstrate uniform EiRP continuity over the blind spot region, which can also be extrapolated to the original DUT.

The results below show the radiated power where the unit is flipped, exposing the section that was in the blind spot to the corresponding angle $\theta = 0^\circ$.

EUT at 2.402 GHz: measured radiated power. Polar 1 = H, Polar 2 = V

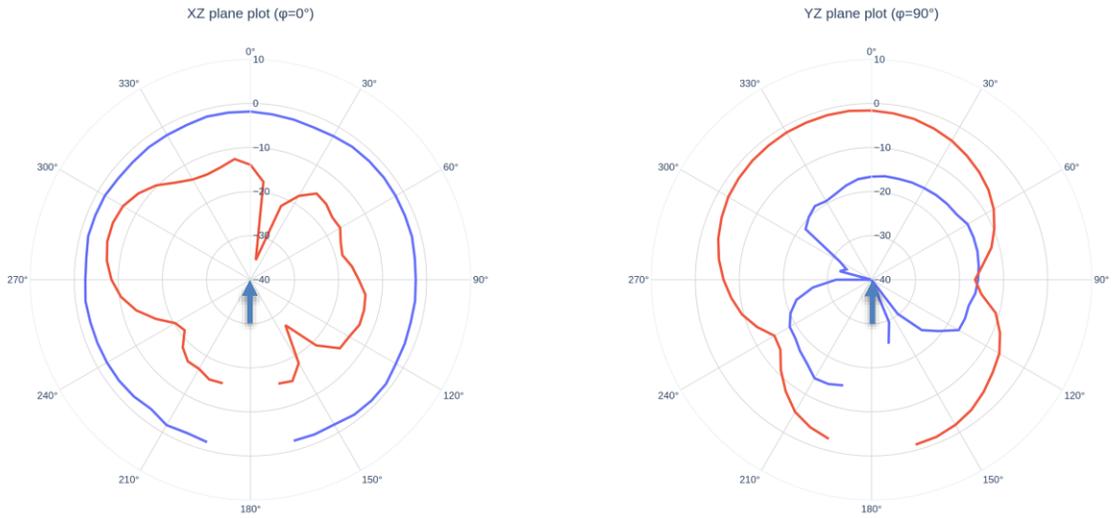


— 2402MHz - Polar 1, TRP=-5.21dBm, Pmax=-3.0dBm, Neb_PB1.h5pmp_009_..._Neb_PB1_X3Q8_Tx_CH02_Pos8dBm_upsidedown.t
— 2402MHz - Polar 2, TRP=-5.21dBm (0.00dBm), Pmax=-11.0dBm, Neb_PB1.h5pmp_009_..._Neb_PB1_X3Q8_Tx_CH02_Pos8dBm_uj
— 2402MHz - Polar 1, TRP=-5.21dBm, Pmax=-15.0dBm, Neb_PB1.h5pmp_009_..._Neb_PB1_X3Q8_Tx_CH02_Pos8dBm_upsidedown.
— 2402MHz - Polar 2, TRP=-5.21dBm (0.00dBm), Pmax=-2.9dBm, Neb_PB1.h5pmp_009_..._Neb_PB1_X3Q8_Tx_CH02_Pos8dBm_up

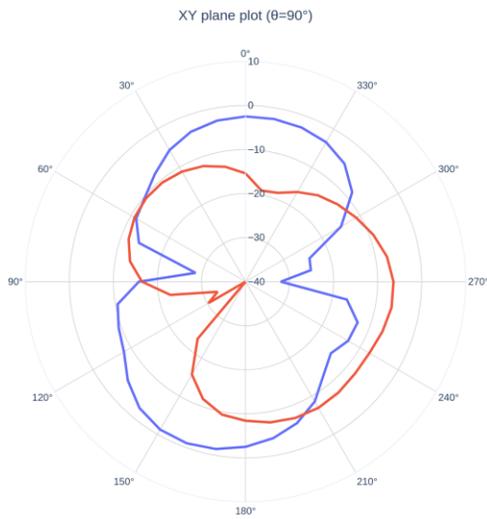


— 2402MHz - Polar 1, TRP=-5.21dBm, Pmax=-2.6dBm, Neb_PB1.h5pmp_009_..._Neb_PB1_X3Q8_Tx_CH02_Pos8dBm_upsidedown.t
— 2402MHz - Polar 2, TRP=-5.21dBm (0.00dBm), Pmax=-9.5dBm, Neb_PB1.h5pmp_009_..._Neb_PB1_X3Q8_Tx_CH02_Pos8dBm_up

EUT at 2.440 GHz: measured radiated power. Polar 1 = H, Polar 2 = V

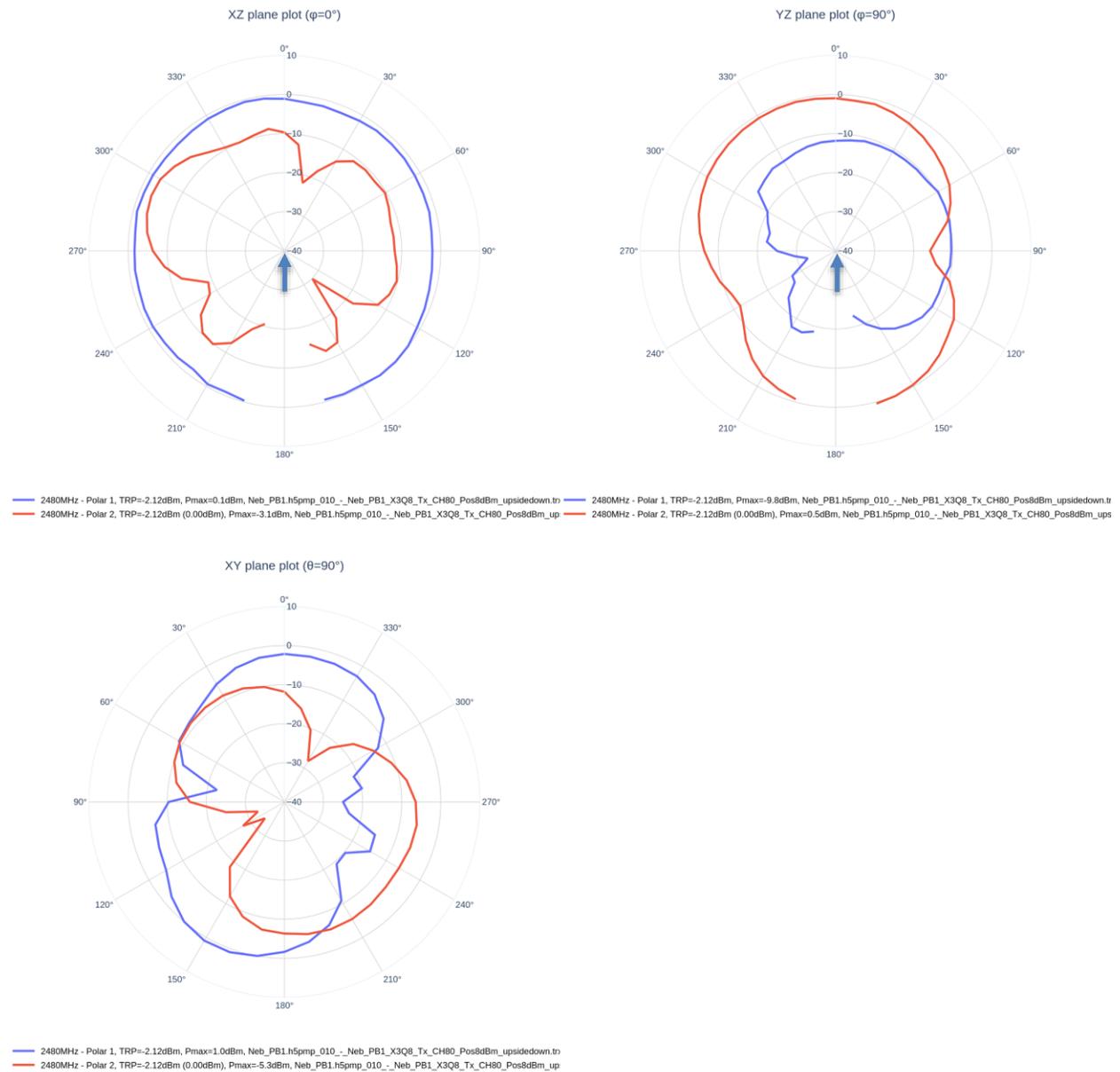


— 2440MHz - Polar 1, TRP=-3.45dBm, Pmax=-1.2dBm, Neb_PB1.h5pmp_007_-_Neb_PB1_X3Q8_Tx_CH40_Pos8dBm_upsidedown.tr
— 2440MHz - Polar 2, TRP=-3.45dBm (0.00dBm), Pmax=-6.2dBm, Neb_PB1.h5pmp_007_-_Neb_PB1_X3Q8_Tx_CH40_Pos8dBm_up
— 2440MHz - Polar 1, TRP=-3.45dBm, Pmax=-14.2dBm, Neb_PB1.h5pmp_007_-_Neb_PB1_X3Q8_Tx_CH40_Pos8dBm_upsidedown.
— 2440MHz - Polar 2, TRP=-3.45dBm (0.00dBm), Pmax=-1.2dBm, Neb_PB1.h5pmp_007_-_Neb_PB1_X3Q8_Tx_CH40_Pos8dBm_up



— 2440MHz - Polar 1, TRP=-3.45dBm, Pmax=-1.0dBm, Neb_PB1.h5pmp_007_-_Neb_PB1_X3Q8_Tx_CH40_Pos8dBm_upsidedown.tr
— 2440MHz - Polar 2, TRP=-3.45dBm (0.00dBm), Pmax=-6.3dBm, Neb_PB1.h5pmp_007_-_Neb_PB1_X3Q8_Tx_CH40_Pos8dBm_up

EUT at 2.480 GHz: measured radiated power. Polar 1 = H, Polar 2 = V



5.0 Conclusion

The additional measurements conducted by flipping the DUT within the Starlab system effectively addressed the blind spot in the theta angle range (165° to 195°). Despite using a different DUT due to technical constraints, the results confirmed consistent EiRP continuity within this region. Furthermore, no unexpected gain was observed in the blind spot, validating the efficiency of the measurement methodology in ensuring accurate radiated power analysis.